

# Spaceborne Aerosol and Ozone Lidars for Air Quality Applications

Rich Ferrare Chris Hostetler Ed Browell John Hair NASA Langley Research Center

Detlef Müller Institute for Tropospheric Research, Leipzig

> David Diner Jet Propulsion Laboratory

NCAR Community Workshop on Air Quality Remote Sensing from Space, February, 2006





- Science Objectives
- Aerosols
  - Current capabilities of spaceborne aerosol lidars
  - High Spectral Resolution Lidar (HSRL)
  - Multiwavelength
     (``3β+2α'') Aerosol
     Retrievals
- Ozone
  - Differential Absorption Lidar (DIAL) Technique
  - Heritage and path to space
- Summary





#### Understanding Global Atmospheric Composition and Predicting Future Evolution

#### Specific Science Questions:

- What are impacts of natural and anthropogenic aerosols on air pollution?
- What are key local and transported aerosol sources?
- How can we improve our ability to model aerosol-mediated heterogeneous chemistry that impacts O<sub>3</sub>, OH, etc. ?
- What is global distribution of tropospheric ozone and how does it change seasonally and interannually?
- What is the relative contribution of photochemical and dynamical processes in determining the distribution of tropospheric ozone?
- What is the impact of ozone on global tropospheric chemistry?

**Aerosol – Current Capabilities and Limitations** 



- + GLAS, CALIOP provide vertical distribution of aerosol
  - Layer heights via backscatter profiles
  - Extinction profiles via inversion
- + CALIOP
  - Extinction profiles constrained with A-train (e.g. MODIS) data
  - Aerosol type can be inferred from backscatter color ratio, depolarization
  - Altitude and back-trajectories to sources also give clues to composition
- Calibration
  - Must calibrate in some region assumed to be free of aerosols and clouds; CALIPSO calibrates at 30-34 km.
  - Can limit calibration to night only; calibration may drift during day
- Measures attenuated total backscatter
  - Aerosol backscatter retrieval requires extinction-to-backscatter ratio S<sub>a</sub>
    - Depends on aerosol composition, size, and shape which are variable
    - Uncertainty in profile of Sa raises potential for structural error in backscatter lidar retrieval

# **High Spectral Resolution Lidar (HSRL)**

# HSRL relies on spectral separation of aerosol and molecular backscatter in lidar receiver.

- HSRL independently measures aerosol and molecular backscatter
  - Can be internally calibrated
  - No correction for extinction required to derive backscatter profiles
  - More accurate aerosol layer top/base heights
- HSRL enables independent estimates of aerosol backscatter and extinction
  - Extinction and backscatter estimates require no  $S_a$  assumptions
  - Provide *intensive* optical data from which to infer aerosol type
  - Measurements of extinction at 2 wavelengths and backscatter at 3 wavelengths enables retrieval of aerosol microphysical parameters and concentration

#### Atmospheric Scattering



#### **Effect of lodine Vapor Notch Filter**





# **Next Step: "3\beta+2\alpha" HSRL retrievals**



- Fundamental data products
  - Backscatter at 3 wavelengths  $(3\beta)$  : 355, 532, 1064 nm 120 m vert., 20 km horiz
  - Extinction at 2 wavelengths (2 $\alpha$ ) : 355, 532 nm 900 m vert., 20 km horiz.
    - Depolarization at 355, 532, and 1064 (dust and contrails/cirrus applications)
- Retrieved, layer-resolved, aerosol microphysical/macrophysical parameters (Müller et al., 1999, 2000, 2001; Veselovskii et al.,2002,2004)
  - Effective and mean particle radius (errors < 30-50%)</li>
  - Concentration (volume, surface) (errors < 50%)</li>
  - Complex index of refraction
    - real (±0.05 to 0.1)
    - imaginary (order of magnitude if < 0.01; <50% if > 0.01)
  - Single scatter albedo (±0.05; error increases for  $r_{eff}$  > 0.3  $\mu$ m)
  - Microphysical retrieval issues
    - Constraining assumptions: positivity, smoothness of size distribution, consistency between retrieved parameters and input optical data
    - Assumes wavelength independent and size independent refractive index
    - Assumes spherical particles; upgrade to spheroids is planned
    - Retrieval is restricted to particle radii > 50 nm
    - Not operational: requires extensive computation time and expert operator, software package in alpha version has been developed for more general use

## "3β+2α" Example microphysical retrieval









- Aerosol layer heights
- Qualitative vertical distribution (backscatter profile)
- Aerosol type vs. altitude
- Extinction profile from backscatter
- Extinction profile with column constraint
- Fine-coarse mode fraction vs. altitude
- Extinction profile
- Complex refractive index vs. altitude
- Aerosol size vs. altitude
- SSA vs. altitude
- Concentration vs. altitude

# **HSRL Technology maturity**



Transmitter

- In general, higher average power is required for HSRL technique over current backscatter lidars (e.g., ~5x over CALIPSO, GLAS)
  - Injection-seeded Nd:YAG laser capable of up to 50 W can be built using current technology
  - Space qualified seed lasers are currently available
  - Required spectral purity and frequency agility have been demonstrated
- Frequency doubling to 532 nm has been demonstrated
- Frequency tripling to 355 nm must be assessed
  - Optical damage due to contamination poses greater problems in UV

#### Receiver

- Iodine absorption cell (532 nm)
  - Used on ground-based systems for many years
  - Poses no technical problem for space
- Multiple Fabry Perot interferometer (355 nm)
  - Demonstrated in the 1970s (Eloranta)
  - Planned for ESA EarthCare
  - Systematic errors associated with spectral characterization of etalon passbands may be an issue
- Mach Zehnder interferometer
  - Demonstrated for ground-based wind measurements
  - CNES developing Mach Zehnder system
  - Systematic errors associated with spectral characterization of etalon passbands may be an issue

# **Spaceborne Ozone Lidar**



- Simultaneous tropospheric and stratospheric ozone profiles with simultaneous aerosol & cloud backscatter & depolarization profiles.
- Will address key global environmental issues including tropospheric chemistry & dynamics and associated ozone and aerosol production & transport for air quality applications; climate change and radiation budget contributions from ozone, aerosols and clouds; and stratospheric chemistry & dynamics and associated surface UV forcing and weather forecasting.
- Measurement resolutions and accuracy goals:

Ozone -	Trop.:	Night: <u>&lt;</u> 2.5 km x 200 km (10%)	
		Day:	
	Strat.:	<u>≤</u> 1 km x 100 km (10%)	
Aerosols -	Trop.:	60 m x 1 km (10%) at $2\lambda$ Backscatter	
	Strat .:	100 m x 10 km (10%) at $2\lambda \int$ not HSRL	

- Spectral Regions: 305-320 nm with 10-12 nm  $\Delta\lambda$  DIAL with two aerosol/cloud channels ( $\lambda_{off}$  & visible/near IR  $\lambda$ ) & one with depolarization.
- Deployment: Small satellite in polar, low Earth orbit with 3 year life.

# **Ozone Lidar Heritage**





PEMWEST-B Latitudinal Ozone Distribution Over Western Pacific





NASA airborne ozone and aerosol lidar measurements have long (~30 year) heritage of global measurements characterizing spatial and vertical distributions of ozone and aerosols for stratospheric and tropospheric applications



#### Major Technological Challenges:

- Transmitter
  - Wavelengths: on-line: 305-308 nm; off-line: 315-320 nm; aerosol wavelength: visible or near IR
  - High-power: >10 W/wavelength with pulse energies of 10 mJ-1 J at pulse rep rates 1 kHz-10 Hz
  - − Lifetime: ≥3 year
- Receiver
  - Large-effective aperture telescope with area >  $4 \text{ m}^2$
  - High-performance UV filters: T >70% with narrow bandwidth
  - High-efficiency (QE >50%), low noise, photon counting detectors

## **Space-based Ozone & Aerosol Lidar Evolution**

#### Progression to Space

- Current airborne
  - DIAL (ozone)
  - HSRL (aerosol) 🗍 🧎
    - MILAGRO/INTEX B air quality/climate
- Instrument Incubator (IIP)
  - Ozone UAV-based
     Global Ozone Lidar
     Demonstrator (GOLD)
  - Ozone+Aerosol –
     Combined HSRL/Ozone
     DIAL
- Support
  - NASA (HQ, LaRC, GSFC, ESTO, LRR, CALIPSO)
  - DOE ASP



# **Summary**



- Aerosol lidars (current)
  - Measure aerosol layer heights and thickness
  - Permit inference of aerosol type
  - Retrieval uncertainties due to relating aerosol backscatter and extinction
- HSRL (future)
  - Unambiguous measurement of extinction and backscatter
  - Possible to implement 2-extinction, 3-backscatter wavelength system for retrieval of microphysical properties and concentration
  - Only known demonstrated remote sensing method for obtaining vertically resolved information on aerosol microphysical properties
- Ozone
  - DIAL technique provides higher vertical resolution than passive techniques
  - Long heritage from airborne measurements
- Potential future spaceborne systems
  - Global, vertically (HSRL) and horizontally (MSPI) resolved measurements of aerosol optical and microphysical properties – Aerosol Global Interactions Satellite (AEGIS)
  - Global, vertically resolved measurements of tropospheric and stratospheric ozone and aerosol distributions - Ozone Research with Advanced Cooperative Lidar Experiments (ORACLE)





## Backup Slides

NCAR Community Workshop on Air Quality Remote Sensing from Space, February, 2006

# **Disadvantage of backscatter lidar: 1 equation, 2 unknowns**





 $\frac{\sigma_p(r)}{\beta_p(r)} = S_p \quad \longleftarrow \text{Assumption of value for extinction-to-backscatter } (S_p) \text{ ratio required for backscatter lidar retrieval}$ 



Measured Signal on Molecular Scatter (MS) Channel:

$$P_{MS}(r) = \frac{C_{MS}}{r^2} F(r) \beta_m(r) \exp\left\{-2\int_0^r \left[\sigma_m(r') + \frac{\sigma_p(r')}{r}\right] dr'\right\}$$
Particulate Extinction

Measured Signal on Total Scatter (TS) Channel:



# **Extinction-to-backscatter ratio variability**



- Multiyear Raman lidar measurements over DOE ARM SGP site found large variations in vertical profile of S<sub>a</sub> occurred 30% of time
- Significant variability in particle size, composition, and/or shape often occurs
- Uncertainty in profile of S<sub>a</sub> raises potential for structural error in backscatter lidar retrieval



# **Heritage and Future Prospects**



U. Wisc Eloranta 1977 – …	Operating ground-based systems for decades; first etalon-based system; first 532 nm iodine vapor filter system;
Colo. St She 1983 – 1998	First vapor filter systems, various wavelengths; first demonstration of temperature measurements
NIES - Liu 1997 – 2001	Ground-based system; 532 nm iodine vapor filter technique
DLR 1998 – 2000	First practical aircraft-based system (no longer functional); 532 nm using iodine vapor filter technique
LaRC 2004 –	Developed aircraft-based system 532 nm HSRL (iodine filter), 1064 backscatter, and depolarization at both wavelengths. Funded to 355 nm HSRL channels through IIP (to be completed by 2008).
CNES 2005 ? – …	"LNG" Leandre upgrade; 355 nm HSRL (Mach Zehnder), 1064 backscatter
ATLID/Earthcare 2012 – …	Spaceborne system; etalon-based receiver; 355 nm

# **Example microphysical retrieval #1**





- From Müller et al., Appl. Opt., 2001
- Data from LACE 98 campaign over Lindenberg, Germany
- Microphysical retrieval performed for upper layer (3-6 km) and compared to in situ aircraft measurements

### Ex. #1- Müller et al. (2001) case study using 3-backsatter and 2-extinction wavelengths



Retrieval results compared to in situ measurements for biomass plume.

Darameter	Lidor Patriaval	Aircraft, in situ	
	Liuai Keulevai	<i>r</i> >1.5 nm	r>50 nm
$r_{eff}$ , $\mu_{m}$	$0.27 \pm 0.04$	$0.24 \pm 0.06$	$0.25 \pm 0.07$
Number concentration, cm <sup>-3</sup>	305±120	640±174	271±74
Surface concentration, <sup>µ</sup> m <sup>2</sup> cm <sup>-3</sup>	145 <b>±</b> 8	110±50	95±55
Volume concentration, <sup>µ</sup> m <sup>3</sup> cm <sup>-3</sup>	13±3	9±5	8±5
$m_R$	$1.63 \pm 0.09$	1.56	1.56
$m_I$	$0.048 \pm 0.017$	0.07	0.07
<i>SSA</i> (532 nm)	0.81±0.03	$0.78 \pm 0.02$	$0.79 \pm 0.02$
<i>SSA</i> (355 nm)	$0.76 \pm 0.06$	—	—
$S_a$ (532 nm) sr <sup>-1</sup>	73±4 (75)	_	_
$S_a$ (355 nm) sr <sup>-1</sup>	51±4 (45)	_	_

# **Spaceborne** $3\beta + 2\alpha$ **HSRL**



	CALIPSO	(rough estimates)
Telescope Diameter	1.0 m	1.5 m
Fundamental Resolution	30 – 60 m vertical, 333 m horizontal	30 m vertical, 70 m horizontal
Backscatter Resolution	120 m vertical, 40 km horizontal	120 m vertical, 20 km horizontal (10% error)
Extinction Resolution	(Indirect) 120 m vertical, 40 km horizontal	(Direct) 900 m vertical, 20 km horizontal (15% error)
Power	200 W @ 700 km	400 W @ 400 km, 825 W at 640 km
Mass	172 kg	265 - 325 kg

NCAR Community Workshop on Air Quality Remote Sensing from Space, February, 2006

## **Measurement Requirements**



- Requirements below are minimums we are currently considering and are driven by
  - 15% accuracy on backscatter and extinction for microphysical retrievals
  - Horizontal and vertical resolutions required to capture relevant aerosol features. Will learn more about relevant aerosol scales with launch of CALIPSO.

Parameter	Resolution	Relative Error
Backscatter	Δx < 150 m Δz < 50 km	< 15%
Extinction	Δx < 1 km Δz < 50 km	< 15%

# **Technology Requirements**



- Transmitter
  - SLM, frequency agile Nd:YAG operating at 1064, 532, and 355 nm
    - Average output power > 50W
    - Rep rates 50-200 Hz
      - higher rep rates are acceptable, but puts more stringent requirement on receiver in terms of solar background rejection.
    - High electrical-to-optical efficiency
  - Issues
    - Lifetime
      - Pump diodes
        - o Need quantitative database on lifetime vs. diode drive current: determine how derating drive current from nominal specs increase lifetime.
      - ➤ UV operation
        - o Expect contamination to be a bigger problem in UV than in visible and near IR. Long-term degradation of coatings due to high power UV exposure should be studied. Contamination and contamination control processes should be studied: absorption by trace organic contaminants more of a problem in the UV.

# **Technology Requirements**



- Receiver
  - Interferometric receiver required for 355 nm HSRL measurement (may also be used at 532 if shows merit over iodine vapor filter technique)
    - Spectral resolution  $\sim 1 \text{ GHz}$
    - Photon efficient
    - Good rejection of solar background
    - High stability/calibration accuracy
      - Accurate calibration of throughput vs. wavelength critical to HSRL application
  - Detectors
    - High QE: >50%
    - Low dark noise
    - Gain sufficient to make amplification noise insignificant
    - Low excess noise factor
  - Telescope
    - Large area: > 1.5 m diameter